Ancient Asteroids

An international observing campaign, that aims towards the characterization of dispersed collisional asteroid families in the Main Belt of our Solar System.



HELLENIC REPUBLIC National and Kapodistrian University of Athens ______ EST. 1837 _____









ASTRONOMICKÝ ÚSTAV UK Matematicko ficikalita V Holisavickach z 180 00 praha B

Step A:

Observability of Primordial family members

Ancient Asteroids

Ancient Asteroids is an international observing campaign, that aims towards the characterization of dispersed collisional asteroid families in the Main Beld of our Solar System.

This Web App finds the observable asteroids, which are members of the Primordial family, based on your location, observation period and preferences. The output is a short list of candidate targets that helps to create an observation plan.

For further information or bug reporting, please contact:

dimathanaso@phys.uoa.gr

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Observability of Primordial Asteroids Application

Vour Location	Observation Pariod	Proformana			
Insert your Location Insert your Loc Longitude (deg): Latitude (deg): Altitude (m): TimeZone: Insert Your Location is: Lo Your TimeZone is:	ocation and Timezon 23.783368 37.968561 250 GMT+2 ongitude: 23.783368 deg	Preferences e Latitude: 37.968561 deg -	Altitude: 0.25 km		 The Web Application can be found here: https://mybinder.org/v2/gh/DimAthanaso/AsteroidsObservability/main?urlpath=%2F voila%2Frender%2FOPA App_v3.0.ipynb 1: Insert your Longitude (in decimal degrees). 2: Insert your Latitude (in decimal degrees). 3: Insert your Altitude (in meters). 4: Choose your Time Zone. 5: Press the "Insert" button.
Find the Observable Asteroids					6: See the confirmation message below.

Step B:

Observability of Primordial family members Ancient Asteroids Ancient Asteroids is an international observing campaign, that aims towards the characterization of dispersed collisional asteroid families in the Main Beld of our Solar System. This Web App finds the observable asteroids, which are members of the Primordial family, based on your location, observation period and preferences. The output is a short list of candidate targets that helps to create an observation plan. For further information or bug reporting, please contact: dimathanaso@phys.uoa.gr © 2021 Dimitrios Athanasopoulos Observability of Primordial Asteroids Application Your Location Observation Period Preferences Insert your Observation period Start date (YYYY-MM-DD): 2021-02-15 Stop date (YYYY-MM-DD): 2021-02-18 Insert 1: Select the "Observation Period" tab. Your Observation Period is from 2021-02-15 at 12:00 (UTC) to 2021-02-18 at 12:00 (UTC). **2:** Insert your Start date in the specified format. Find the Observable Asteroids **3:** Insert your End date in the specified format. 4: Press the "Insert" button.

5: See the confirmation message below.

Step C:

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Observability of Primordial Asteroids Application

Your Location Observation Period Preferences							
Insert your instrument limitations and preferences The target max. ap. magnitude (mag): 17 The minimum observation time (hours): 4 Insert 17 You want to observe targets with ≤ 17.0 mag for at least 4.0 hours.	 Select the "<u>Preferences</u>" tab. Insert the maximum apparent magnitude of the target that you can observe (in mag). Insert the minimum observation time (≥4 h) that you can observe (in hours). 						
Find the Observable Asteroids	4: Press the "Insert" button. 5: See the confirmation message below						
	6: Press the "Find the Observable Asteroids" button.						

Wait to search asteroids...

Observability of Primordial family members

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Observability of Primordial Asteroids Application

Your Location	Observation Period	Preferences			
Insert your instru	ment limitations a	and preferences			
The target max. ap. n	nagnitude (mag): 17				
The minimum observat	tion time (hours): 4				
You want to observe tar	gets with ≤ 17.0 mag fo	or at least 4.0 hours.			
	Fin	d the Observable Aster	roids	Initialisation	The Web Application searches for observable asteroids
Searching:					When the progress bar is filled, the results will be
					displayed automatically.

Step D:

Initialisation completed!

The observable asteroids

Target	Max. Duration (h)	Min. V (mag)	P (h)	λ_1 (°)	β ₁ (°)	λ_2 (°)	β ₂ (°)	Priority Rate (%)
4422 Jarre (1942 UA)	7.0	16.506	5.43	-	-	•	-	14.0
2768 Gorky (1972 RX3)	4.0	16.805	4.51	-	-	-	-	14.0
25343 (1999 RA44)	9.0	16.784	0.00	-	-	-	-	11.0
43725 (1978 RK9)	7.0	16.704	0.00	-	-	-	-	11.0
2792 Ponomarev (1977 EY1)	4.0	16.919	137.57	-	-	-	-	10.0
428 Monachia (A897 WA)	8.0	15.281	3.63	-	-	-	-	8.0
2536 Kozyrev (1939 PJ)	5.0	16.514	7.19	-	-	-	-	8.0
853 Nansenia (A916 GN)	7.0	14.917	7.93	-	-	-	-	7.0
783 Nora (A914 FB)	8.0	15.212	34.40	-	-	-	-	1.0
282 Clorinde (A889 BA)	4.0	14.987	6.42	-	-	-	-	1.0
1244 Deira (1932 KE)	7.0	14.401	210.60	107	-56	314	-46	0.0
220 Stephania (A881 KA)	4.0	15.984	18.20	26	-50	223	-62	0.0

The ephimerides were produced by the JPL's HORIZON system. The physical parameters have been retrieved from the Minor Planet Physical Properties Catalogue (Delbo, Walsh, Bolin, Avdeilidou, Morbidelii, 2017, 'Identification of a primordial asteroid family constraints the original planetesimal population', Science, 357,

1028 - 1029) and the spin pole data from the DAMIT (Durech, Sidorin, Kaasalainen, 2010, "DAMIT: a database of asteroid models", A & A, 513, A46).

To characterise an asteroid family concerning the rotational state of its members, a homogeneous observing sample, along the distance from the centre of the family is required. The following diagram shows the priority rate for each Primordial family member as a function of its position from the centre of the family. The crosses indicate the position of the members within the family. The red crosses indicate members with known spin state.



Step E:



Ancient Asteroids - Observer guide: Web Application

Your Observation Plan is ready!

Create the Observation Plan

The diagrams below shows the observation plan for the selected target, based on your location. The table lists the observational parameters for the given location and date, including the duration of the observing window, the coordinates of the target at the beginning and end of the observing window, as well as the time, when the target is high enough above the horizon (airmass >3). Note that the observation time is given in UTC, and the time zone in added next to it. The apparent magnitude of the target is changing as a function of time, as shown on the bottom diagram.

Location: (37.968561,23.783368) Timezone: Etc/GMT-2

Mid-Date	Duration (h)	V (mag)	Observation Start (UTC + Timezone)	Observation End (UTC + Timezone)	Start Point [RA,DEC]	End Point [RA,DEC]	Solar Phase (°)	PABLon (°)	PABLat (°)
2021-02-16	7.833333	16.529	2021-02-15 19:50:00+02:00	2021-02-16 03:40:00+02:00	[174.29718, 4.37285]	[174.23634, 4.40998]	9.5190	168.2477	1.4482
2021-02-17	8.000000	16.505	2021-02-16 19:40:00+02:00	2021-02-17 03:40:00+02:00	[174.11631, 4.48646]	[174.05248, 4.52508]	9.0947	168.2478	1.4731
2021-02-18	7.833333	16.481	2021-02-17 19:40:00+02:00	2021-02-18 03:30:00+02:00	[173.92911, 4.6029]	[173.86497, 4.64135]	8.6684	168.2453	1.4978

The target 4422 Jarre (1942 UA) is observable for these dates:



Now you are ready to observe!

That's it!

You can still change the selected target and start over again.

Contact the Team

Any questions, queries or other communication, related to the <u>Ancient Asteroids</u> project, should be addressed to:

Kosmas Gazeas

PI and Scientific Coordinator, National and Kapodistrian University of Athens kgaze@physics.auth.gr, kgaze@phys.uoa.gr

AND COPY to:

Dimitrios Athanasopoulos

PhD candidate, National and Kapodistrian University of Athens <u>dimathanaso@phys.uoa.gr</u>

The email should have the title: "Ancient Asteroids", otherwise it will be omitted.

http://users.uoa.gr/~kgaze/research_asteroids_en.html