



# Establishing and Using the real-time Neutron Monitor Database (NMDB)

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# Outline

## ➤ Introduction

- ✓ Sun, Earth & Cosmic Ray Connection
- ✓ Cosmic Ray Travel
- ✓ Cosmic Ray Spectrum
- ✓ Cosmic Ray Shower
- ✓ Neutron Monitor
- ✓ Cosmic Ray Modulation
- ✓ Neutron Monitors Distribution

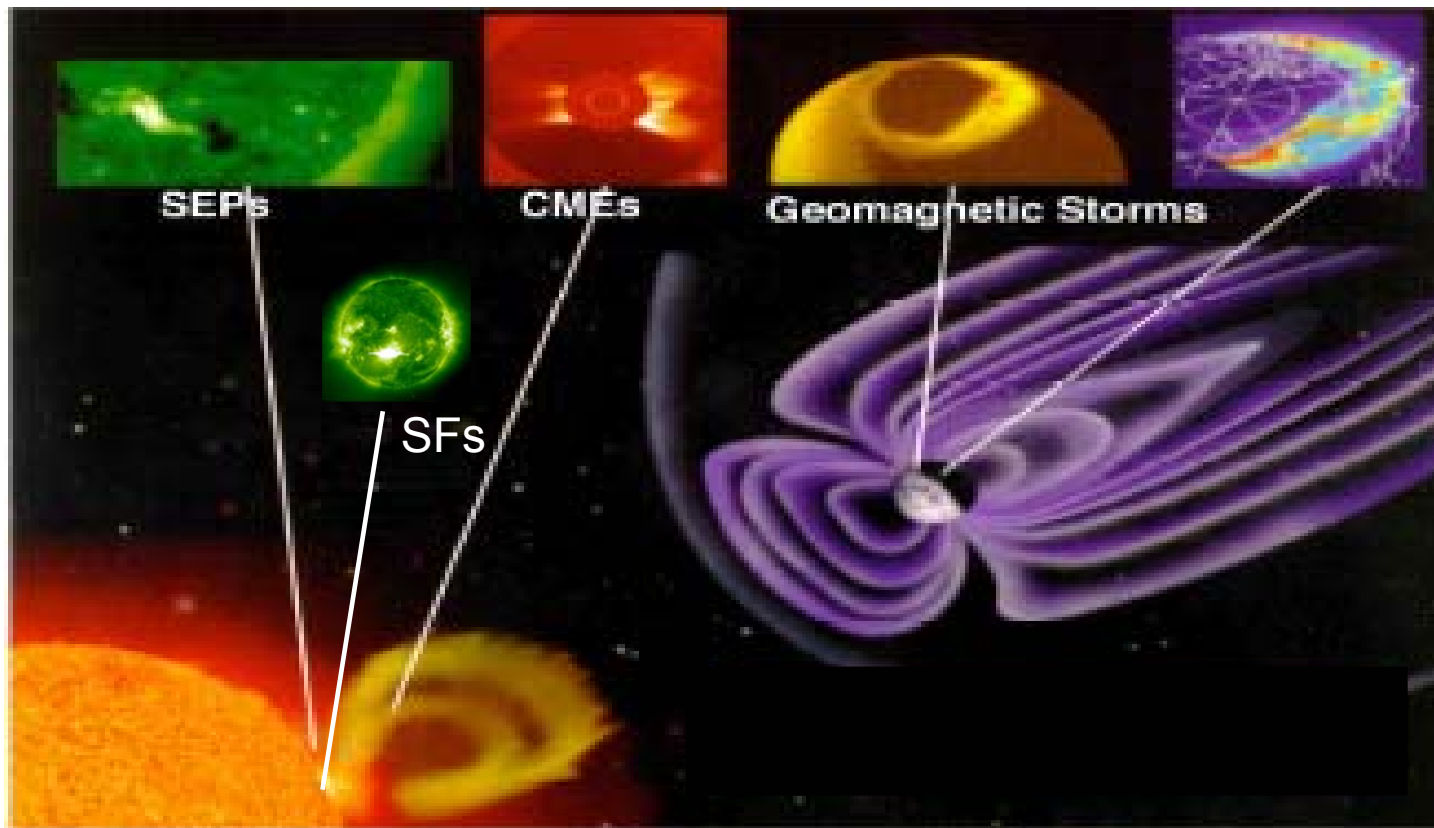
## ➤ NMDB

- ✓ Background of the project
- ✓ Structure of NMDB
- ✓ Goals of the project
- ✓ Creating the database
- ✓ Using NMDB / User tools
- ✓ Applications with 1-min real-time data

## ➤ Conclusions



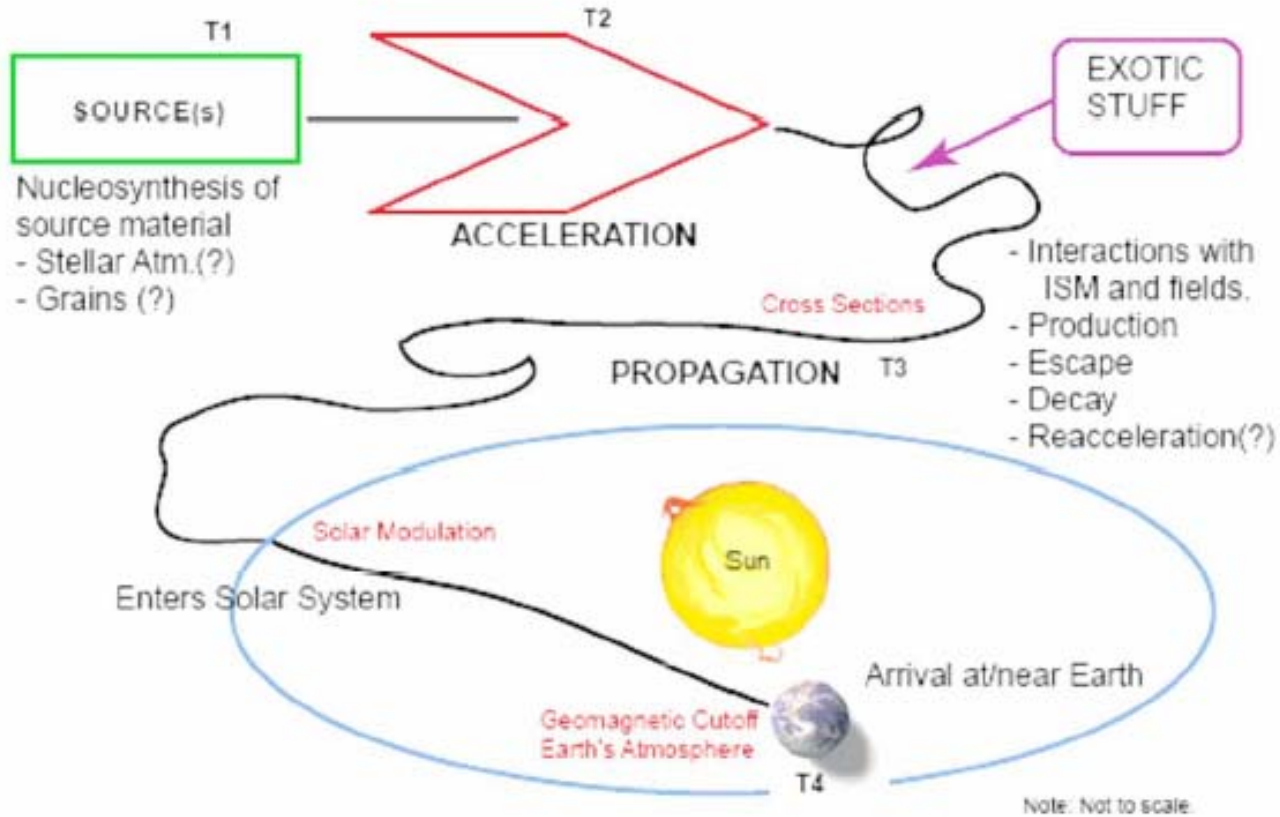
# Sun-Earth & Cosmic Ray Connection



The Sun produces energetic particles and cosmic rays and modulates the galactic cosmic ray flux



# Cosmic Ray Travel





# Integral Spectrum of Cosmic Rays

## Primary Cosmic Rays

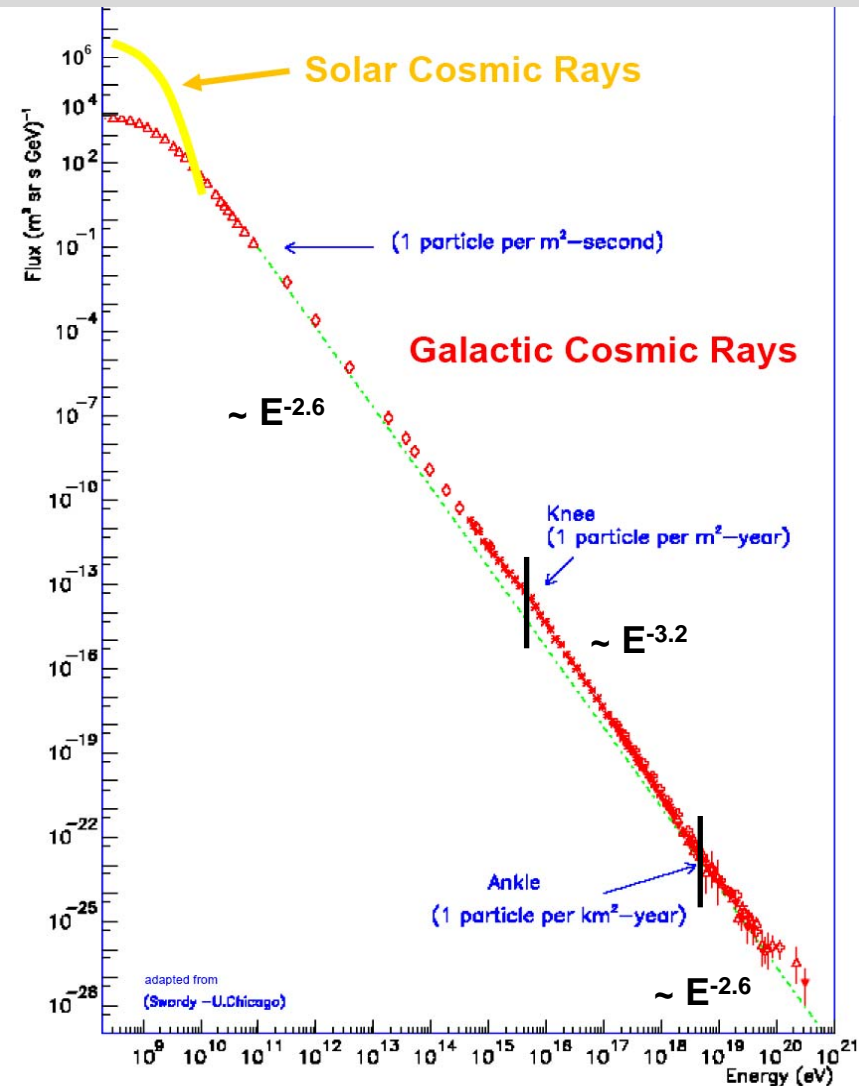
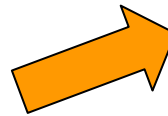
- 90 % protons
- 9 %  $\alpha$  particles
- 1 % heavier nuclei

Energies:  $10^6 - 10^{21}$  eV

$10^9 - 10^{15}$  eV:  $\gamma \sim 2,6$

$10^{15} - 10^{18}$  eV:  $\gamma \sim 3,2$

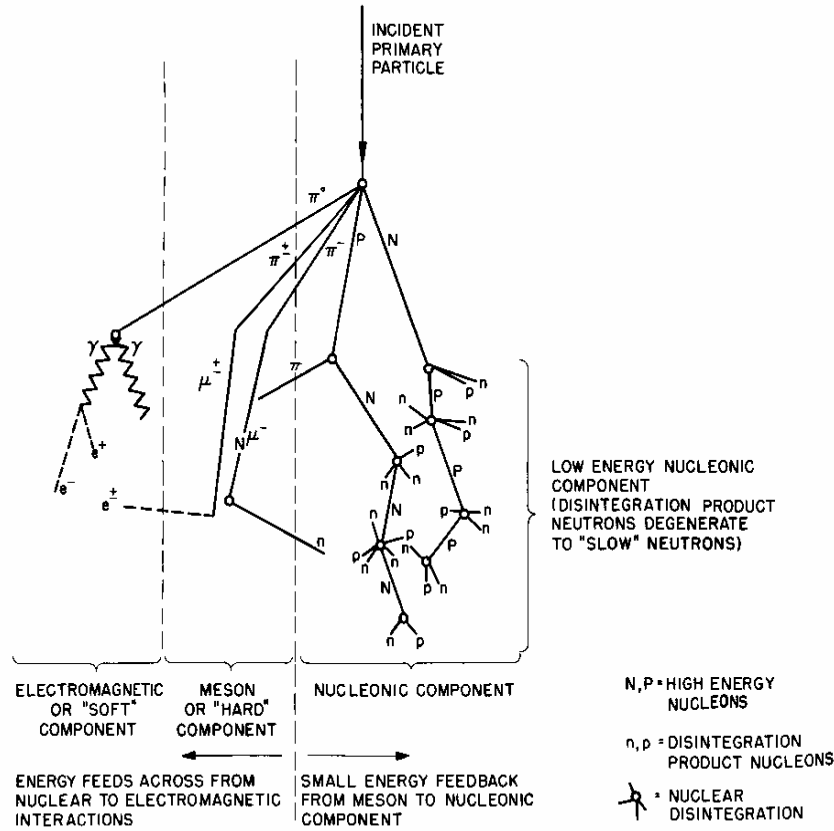
$> 10^{18}$  eV:  $\gamma \sim 2,6$



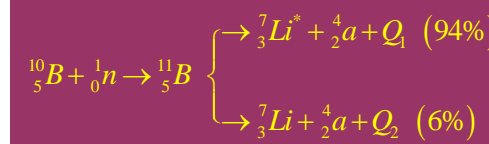
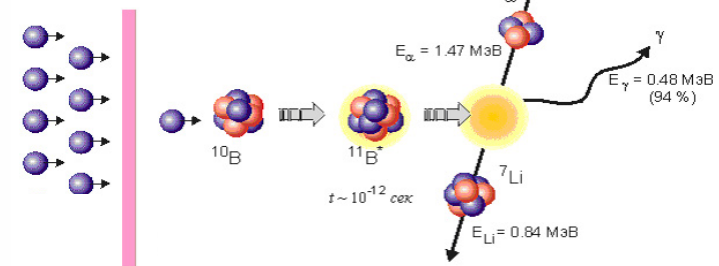


# Cosmic Ray Cascade / Neutron Monitors

✓ **Neutron monitors** are standard devices located at different points on the globe.

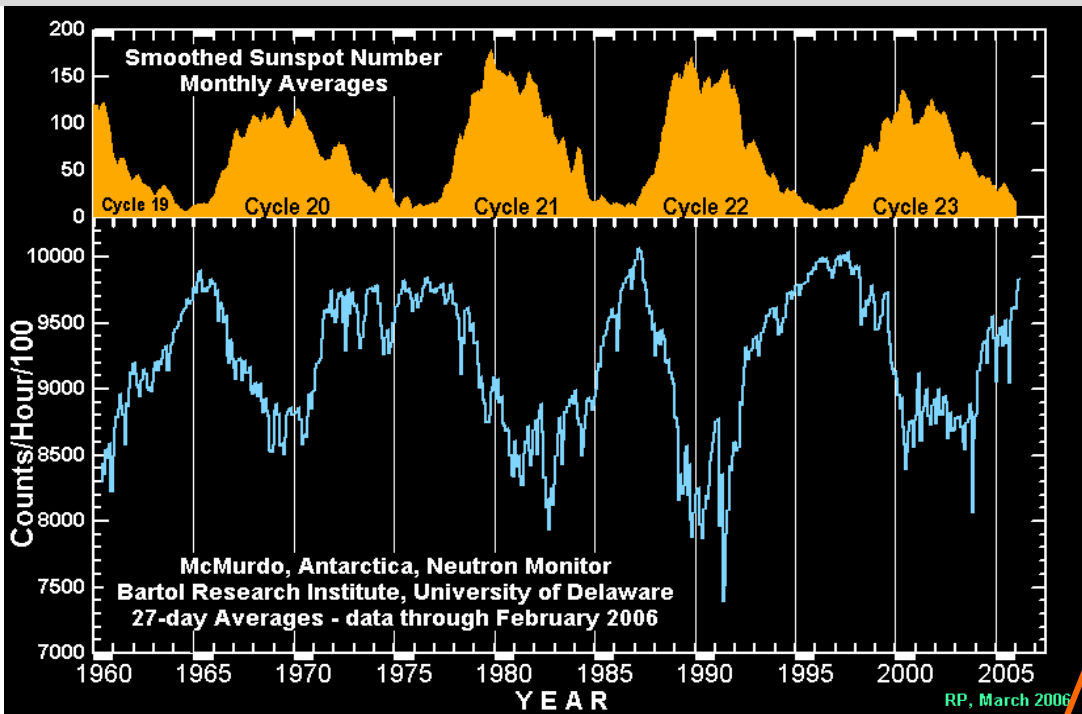


Schematic Diagram of Cosmic Ray Shower



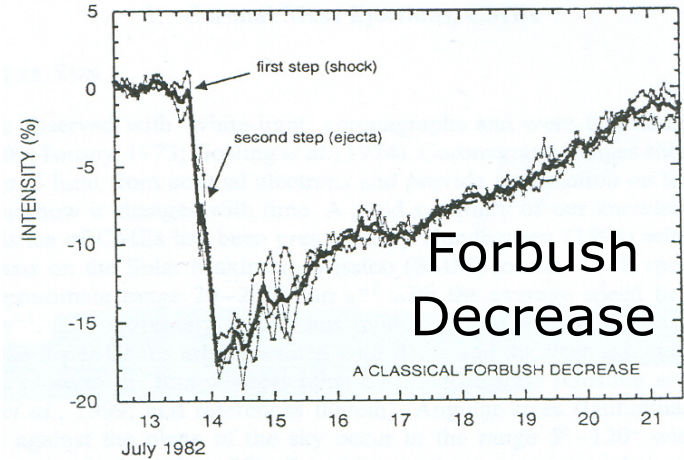


# Cosmic Ray Modulation

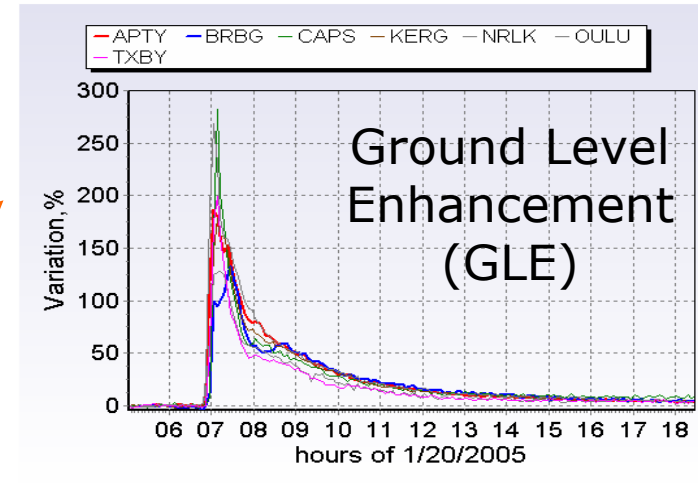


Modulation of Cosmic Rays

Heliospheric phenomena registered at Earth



Forbush 1957; Belov 2001

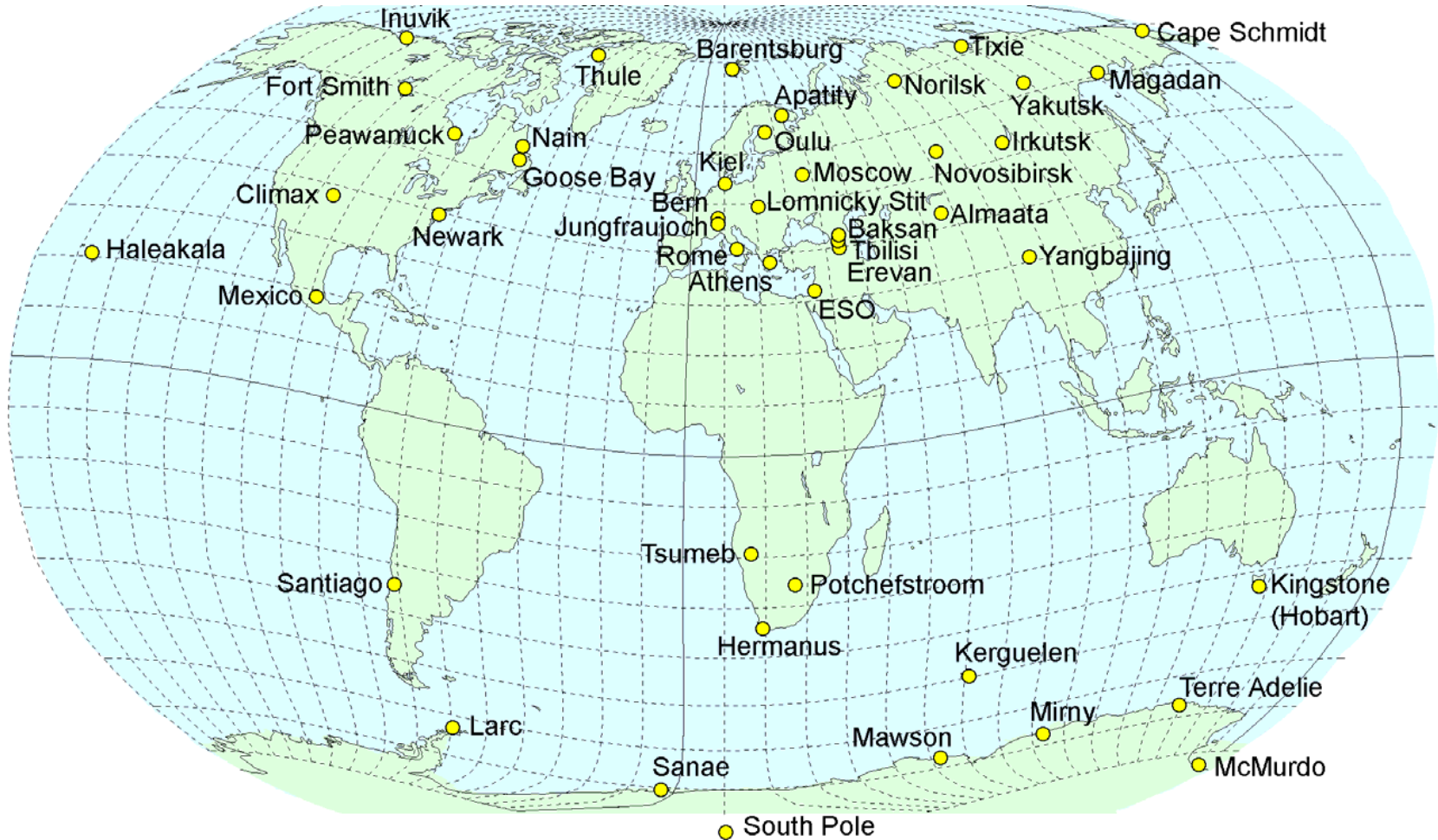


Bonrie et al. 2001





# Neutron Monitor Distribution







# Real-time Neutron Monitor Database (NMDB)

<http://www.nmdb.eu>



- 12 partners
- 28 NM stations in real-time





# Background of the project

## NM Advantages

- ◆ Cosmic ray intensity has been continuously measured with Neutron monitors since the International Geophysical Year 1957/58
- ◆ Approximately 50 Neutron monitors are operating worldwide
- ◆ Neutron monitors measurements can not be scrambled by any intense event
- ◆ Cost-effective reliable detectors
- ◆ Wide detection range
- ◆ Ready to use

## What did we need to take care of

- ◆ High resolution data (1-min or better) are needed to study solar cosmic ray events
- ◆ Only few NM stations could provide data in real-time
- ◆ No common data format, no data center for high resolution data in real time
- ◆ No common real-time applications



# NMDB structure

NMDB

WP1

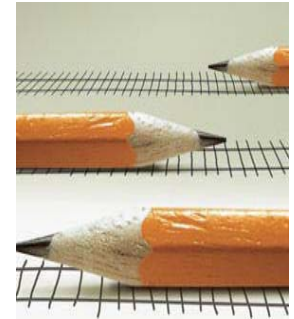
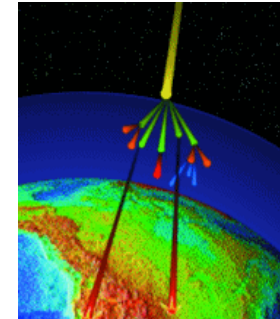
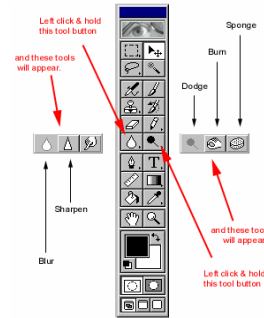
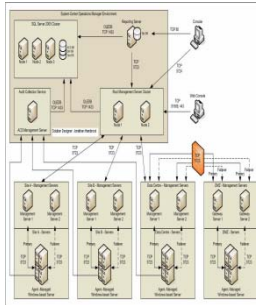
WP2

WP3

WP4

WP5

WP6



WP1 leader CAU

WP2 leader CAU / ALCALA

WP3 leader NKUA

WP4 leader OULU

WP5 leader IZMIRAN

WP6 leader OBSPARIS

Management

Architectural Design

Data Acquisition

User Tools

Applications

Education/ Outreach



# Goals of the project

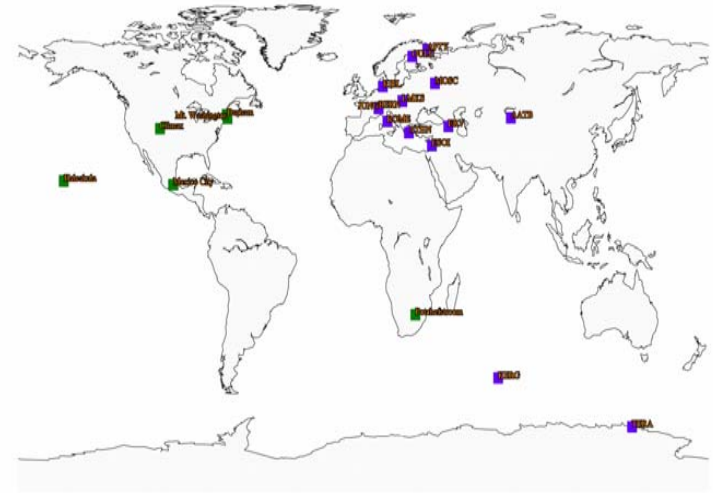
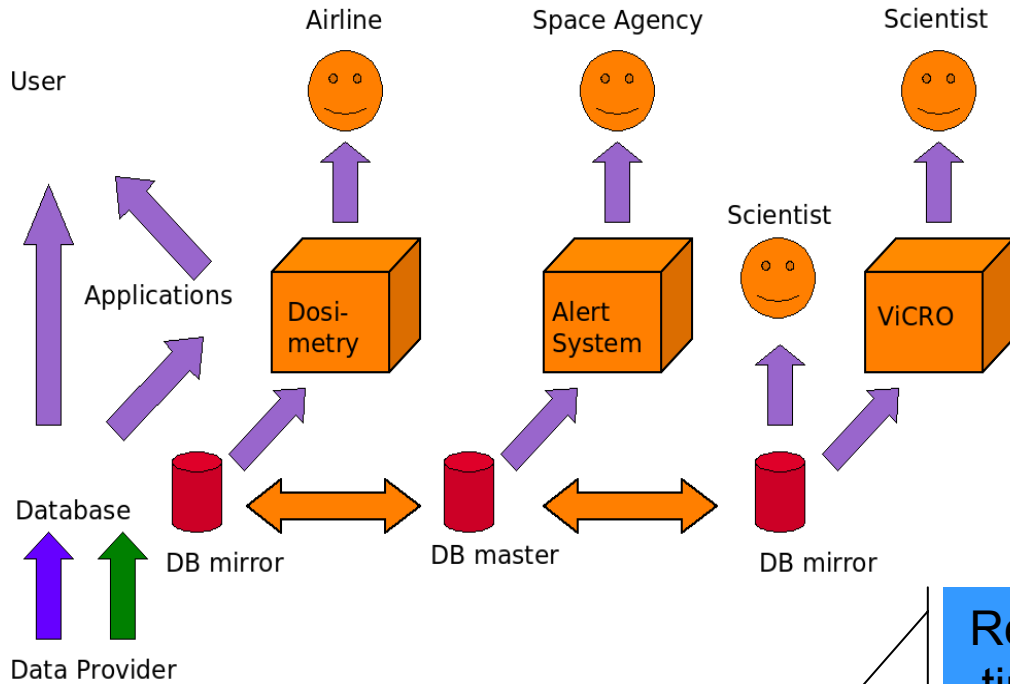
- ◆ Upgrades stations and connections to make high-resolution NM data available in real-time
- ◆ Collect high resolution NM data in real-time, make it public available in common format
- ◆ Build a state of the art, centralized database with distributed mirrors
- ◆ Create an innovate registration system for cosmic rays
- ◆ Develop application tools using NM data
- ◆ Create a public outreach website to inform about cosmic rays and possible effects on humans, technological systems and environment
- ◆ Realize the first ever Training Course on Cosmic Ray Physics and Applications

**Easiest Access to NM data**



# Creating the database

## Schematic concept of NMDB



Real-time

Historical

Meta-data

1 min NM data  
1 hour NM data  
environmental data

1 min NM data  
1 hour NM data  
environmental data

Application's meta-data



# Using NMDB

Station  Table

## Checking NMDB Stations at a Glance

	Station	Table	<a href="#">Original data (1min)</a>	<a href="#">1 hour data</a>	<a href="#">Environmental data</a>
1.	<a href="#">Almaaty</a>	AATB	✔ Online	✔ Online	✘ Offline
2.	<a href="#">Apatity</a>	APTY	✔ Online	✔ Online	✘ Offline
3.	<a href="#">Aragats</a>	ARNM	✔ Online	✘ Offline	✘ Offline
4.	<a href="#">Athens</a>	ATHN	✔ Online	✔ Online	✘ Offline
5.	<a href="#">BKSN</a>	BKSN	✔ Online	✔ Online	✘ Offline
6.	<a href="#">ERV3</a>	ERV3	✘ Offline	✘ Offline	✘ Offline
7.	<a href="#">ERV</a>	ERV	✘ Offline	✘ Offline	✘ Offline
8.	<a href="#">Mt Hermon</a>	ESOI	✘ Offline	✘ Offline	✘ Offline
9.	<a href="#">IRK2</a>	IRK2	✘ Offline	✘ Offline	✘ Offline
10.	<a href="#">IRKT</a>	IRKT	✘ Offline	✘ Offline	✘ Offline
11.	<a href="#">Jungfrauoch, IGY</a>	JUNG1	✔ Online	✔ Online	✔ Online
12.	<a href="#">Jungfrauoch, NM64</a>	JUNG	✔ Online	✔ Online	✔ Online
13.	<a href="#">Kerguelen</a>	KERG	✘ Offline	✘ Offline	✘ Offline
14.	<a href="#">Kiel</a>	KIEL	✘ Offline	✘ Offline	✘ Offline
15.	<a href="#">LEIK</a>	LEIK	✘ Offline	✘ Offline	✘ Offline
16.	<a href="#">Lomnický štít</a>	LMKS	✔ Online	✔ Online	✔ Online
17.	<a href="#">Mobile Cr lab.</a>	MCRL	✔ Online	✔ Online	✘ Offline
18.	<a href="#">Magadan</a>	MGDN	✘ Offline	✔ Online	✘ Offline
19.	<a href="#">Moscow</a>	MOSC	✘ Offline	✘ Offline	✘ Offline
20.	<a href="#">MRNY</a>	MRNY	✔ Online	✘ Offline	✘ Offline
21.	<a href="#">Nor-Amberd</a>	NANM	✔ Online	✘ Offline	✘ Offline
22.	<a href="#">Norilsk</a>	NRLK	✘ Offline	✘ Offline	✘ Offline
23.	<a href="#">Novosibirsk</a>	NVBK	✘ Offline	✘ Offline	✘ Offline
24.	<a href="#">Oulu</a>	OULU	✔ Online	✔ Online	✘ Offline
25.	<a href="#">Rome</a>	ROME	✔ Online	✔ Online	✘ Offline
26.	<a href="#">Terre Adelie</a>	TERA	✘ Offline	✘ Offline	✘ Offline
27.	<a href="#">TXBY</a>	TXBY	✘ Offline	✘ Offline	✘ Offline
28.	<a href="#">Yakutsk</a>	YKTK	✘ Offline	✘ Offline	✘ Offline
			Online: 13 Offline: 15	Online: 11 Offline: 17	Online: 3 Offline: 25

1 – min data

1 – hour data

environmental data





# Using NMDB

## NMDB Event Search Tool - NEST

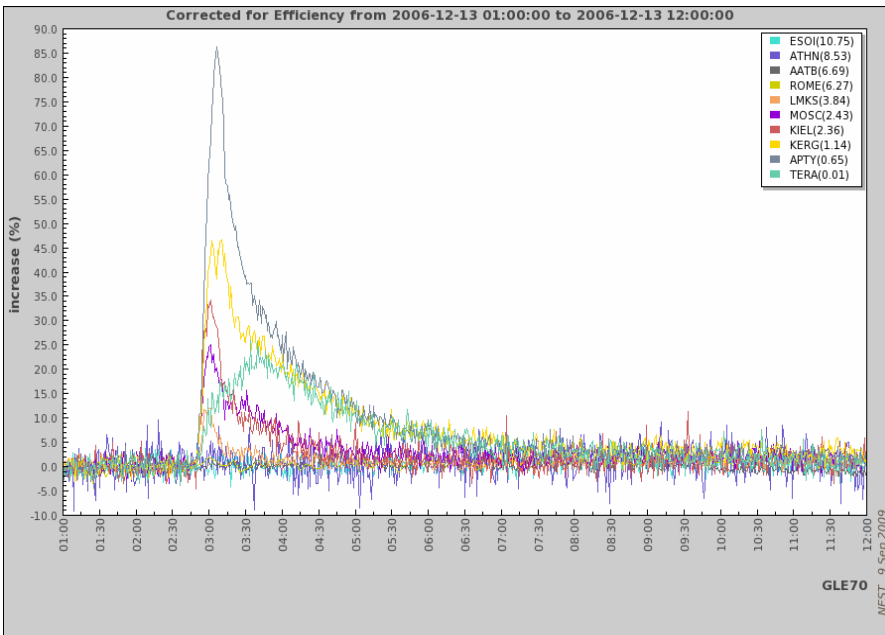
Choose the stations you want

Check the latest submitted data to NMDB – See if everything works !

Take data & plots

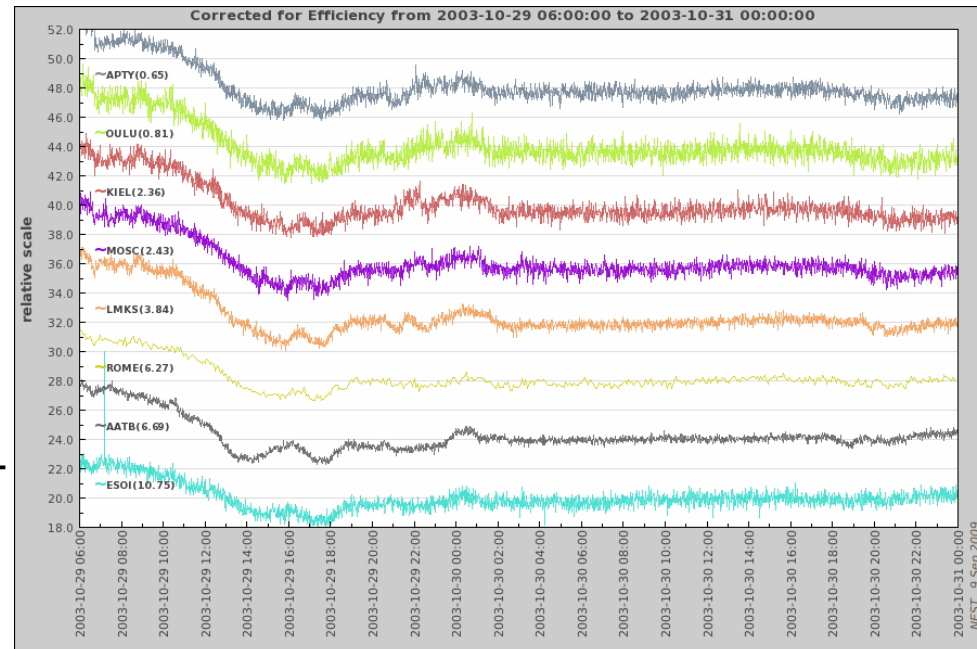


# Using NMDB



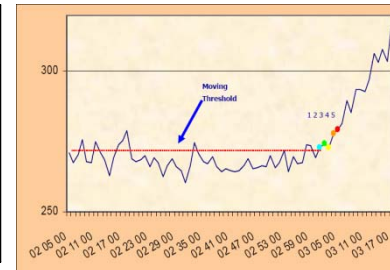
✓ GLE illustrated by NEST

✓ Forbush decrease illustrated by NEST

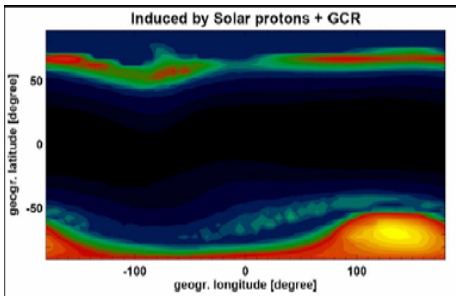




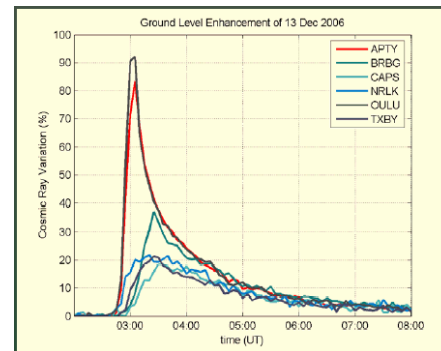
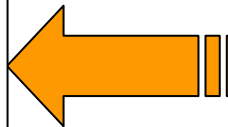
# Applications with 1-min real-time data



In real or quasi-real time



Atmosphere's ionization



GLE Model approach



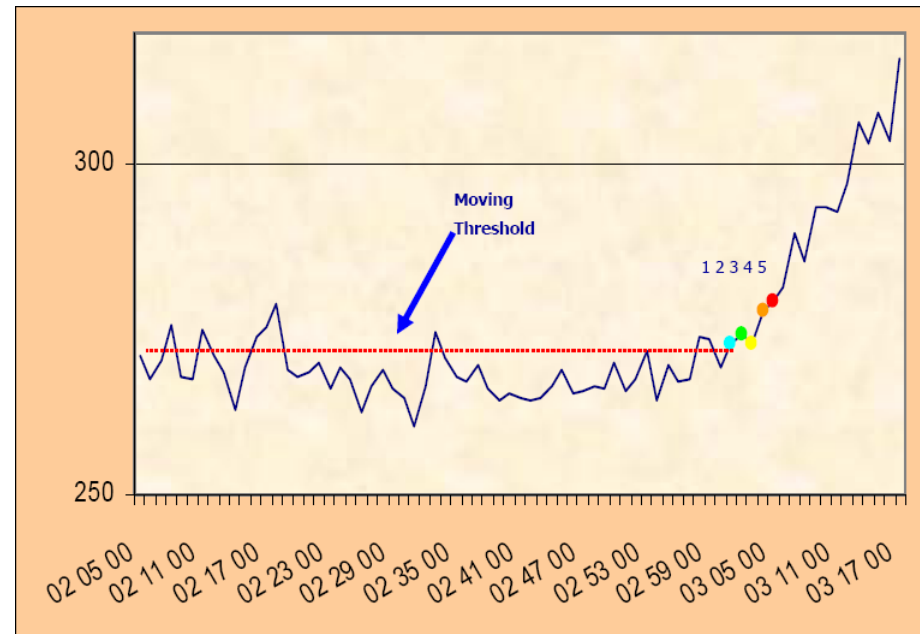
# Applications with 1-min real-time data

## • GLE Alert Algorithm

• Several groups (NKUA, IZMIRAN, TAU, ALMATY), participating at NMDB project, hold various GLE Alert functions – some of which operate in real-time. Under the cooperation of NMDB, these groups will work together and will provide the best possible Alert

### ➤ Steps of the Station Alert

1. We define a moving threshold
2. When the last measurement exceeds this threshold, the algorithm marks a pre-alert point
3. If we get 5 pre-alert points in succession we define a Station Alert





## • GLE Alert Algorithm

### ➤ Steps of the General Alert

1. A Supervision program named program named Check For alert checks every minute the status of every station.
2. If this program detects at least three stations in "station alert mode" then produces a General GLE Alert Signal

Station	Status
AATB	Keep Searching
APTY	Keep Searching
ATHN	Keep Searching
ERV	Keep Searching
ERV3	Keep Searching
ESDI	Keep Searching
IRKT	Keep Searching
JUNG	Keep Searching
JUNG1	Keep Searching
KERG	Keep Searching
KIEL	Keep Searching
LMKS	Keep Searching
MCRL	Keep Searching
MGDN	Keep Searching
MOSC	Keep Searching
MRNY	Keep Searching
NRLK	Keep Searching
NVBK	Keep Searching
OULU	Keep Searching
ROME	Keep Searching
TERA	Keep Searching
YKTK	Keep Searching
TXBY	Keep Searching

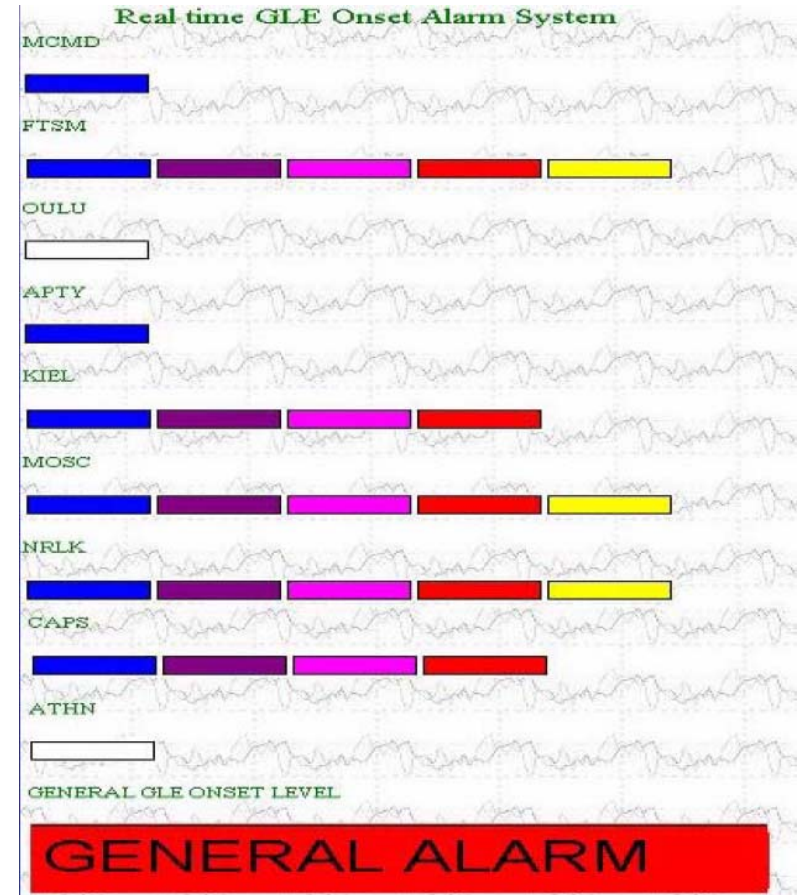
Information Text  
LAST GLE ALERT 2009-02-22 22:14 MCMD FTSM OULU



## • GLE Alert Algorithm

# The first real-time GLE Alert signal (NKUA, 2009)

- ✓ The Ground Level Enhancement (GLE) event of the December 13, 2006 – GLE70
- ✓ Using the steps and the algorithms described at the previous slides, a real-time Alert signal was registered
- ✓ This is the reconstructed display of the Alert at the webpage. You can notice that three stations (FTSM, MOSC, NRLK) provide the General Alert stage







# Validation of GLE Algorithm

GLE No	Event date	Flare time (UT)	Flare	GOES Alert(100 MeV, >1 pfu)	NM Stations Alert
60	15 April 2001	13:19	X14.4	14:21	13:59
61	18 April 2001	02:11	C2	03:11	02:43
62	04 Novemeber 2001	16:03	C2	17:07	16:50
63	26 December 2001	04:32	C2	06:14	06:07
64	24 August 2002	00:49	X3.1	01:48	01:35
65	28 October 2003	09:51	X17.2	11:51	11:17
66	29 October 2003	20:37	X10.0	-----	21:08
67	2 November 2003	17:03	X8.3	17:56	17:39
68	17 January 2005	-----	-----	-----	NO GLE
69	20 January 2005	06:36	X7.1	07:04	06:52
70	13 December 2006	02:48	X3.4	03:12	02:53

✓ The above table shows that GLE Alert which depends solely on ground based observations precedes the one triggered by GOES at a time window of **7 - 34 minutes**. There is also a single case (GLE66) where the NM GLE algorithm issues an Alert while satellite data did not.



# GLE Alert Algorithm

## Real-time Illustration of GLE Alert

Online real-time GLE Alert

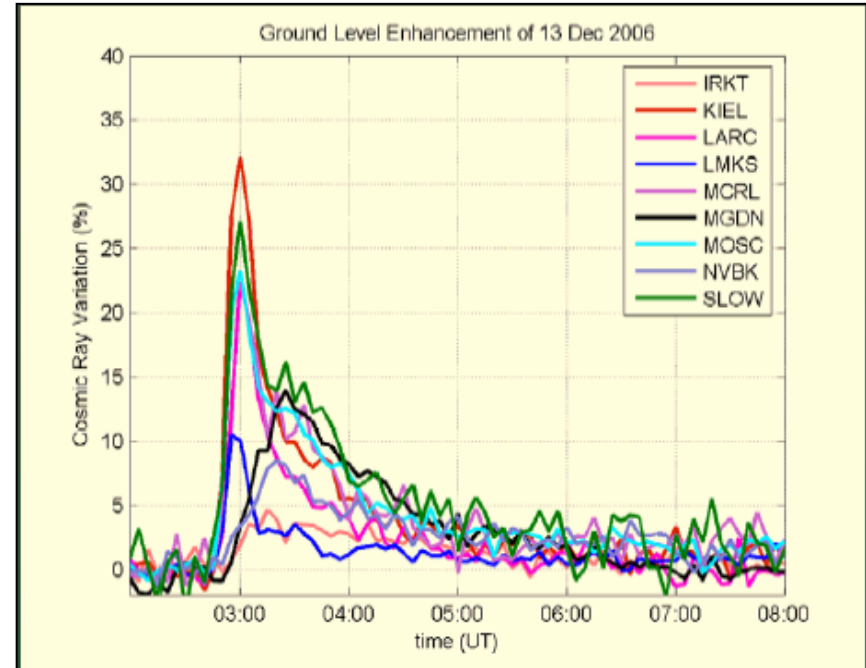
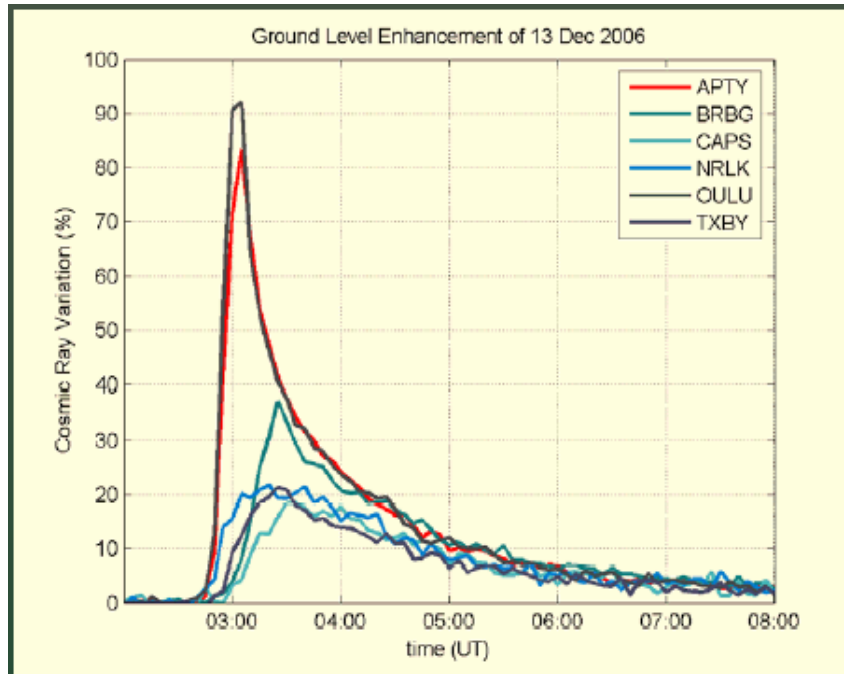
RESULT of the GLE Alert for the last 12 minutes

DateTime	Total Stations	Used Stations	Stations with increase >2 sigma	Watch	Warning!	Alert !
2009-09-16 08:33:00+00	15	8	1			
2009-09-16 08:34:00+00	15	8	1			
2009-09-16 08:35:00+00	15	7	0			
2009-09-16 08:36:00+00	15	7	1			
2009-09-16 08:37:00+00	15	7	2			
2009-09-16 08:38:00+00	15	7	3	✓	✓	
2009-09-16 08:39:00+00	15	7	1			
2009-09-16 08:40:00+00	15	5	1			
2009-09-16 08:41:00+00	15	5	3	✓	✓	
2009-09-16 08:42:00+00	15	5	1			
2009-09-16 08:43:00+00	15	3	0			
2009-09-16 08:44:00+00	15	0	0			



# Applications with 1-min real-time data

## • GLE Modelling – NM BANGLE

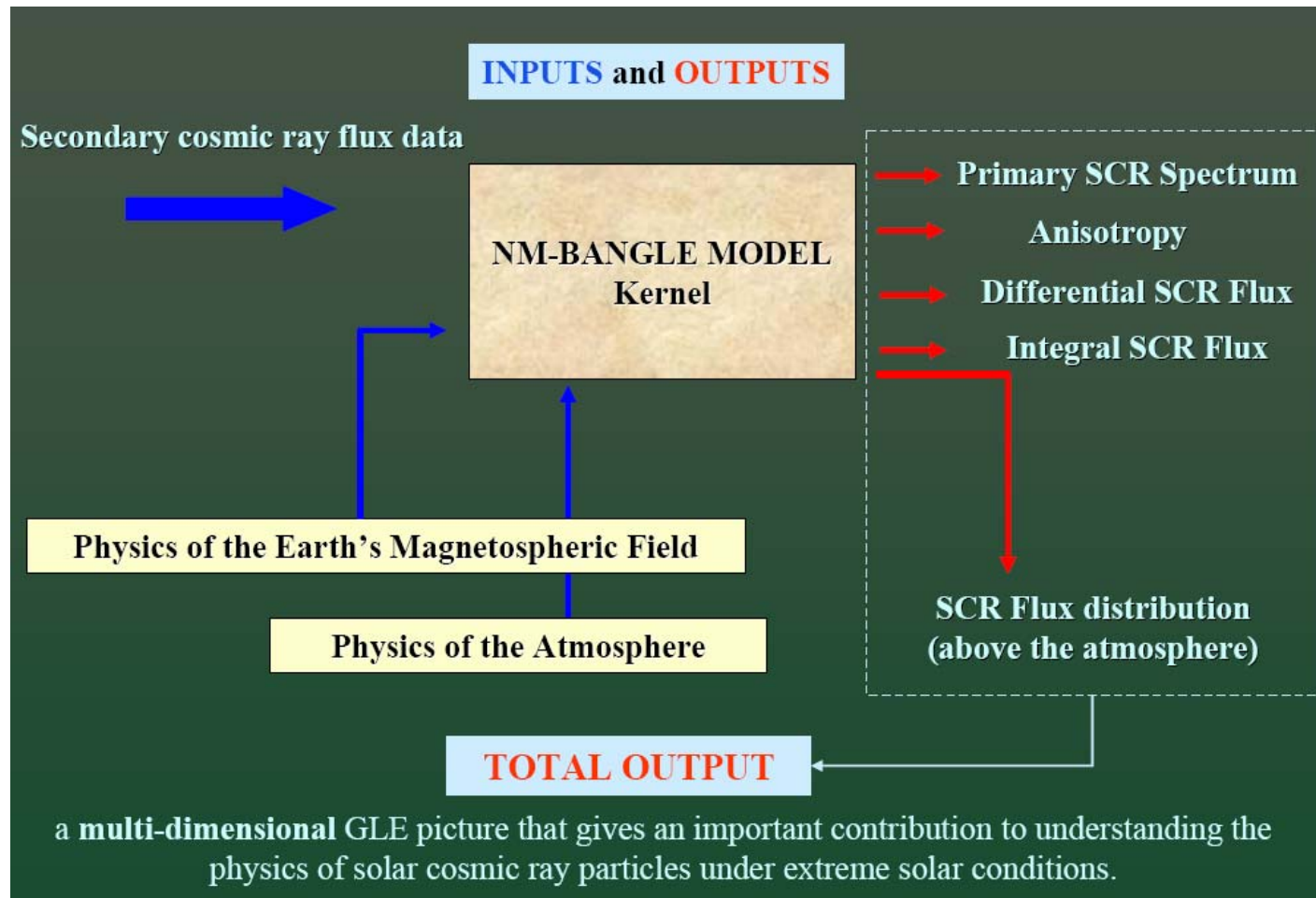


- Several groups participating at NMDB project, created GLE models. One of the most sophisticated ones is the NM-BANGLE



# • GLE Modelling – NM BANGLE

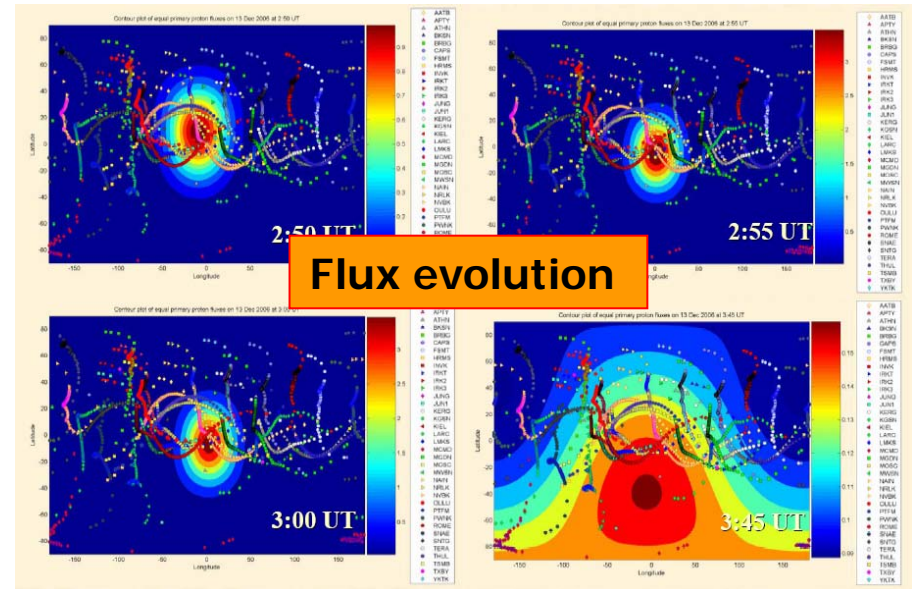
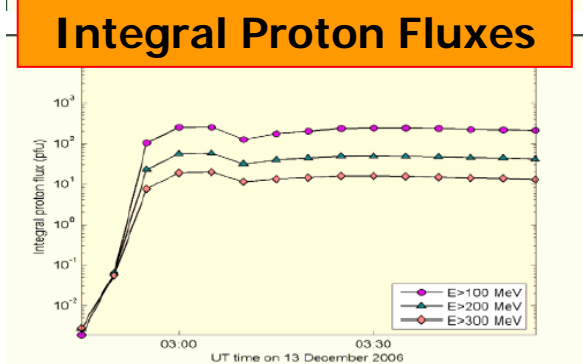
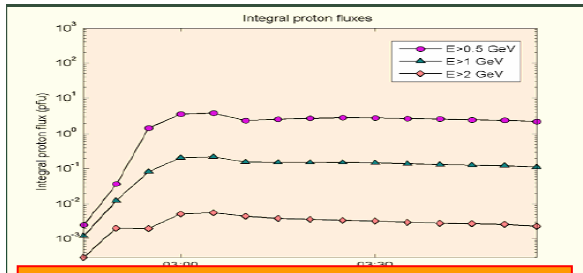
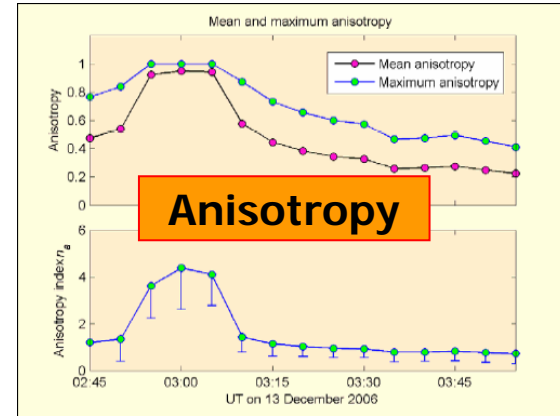
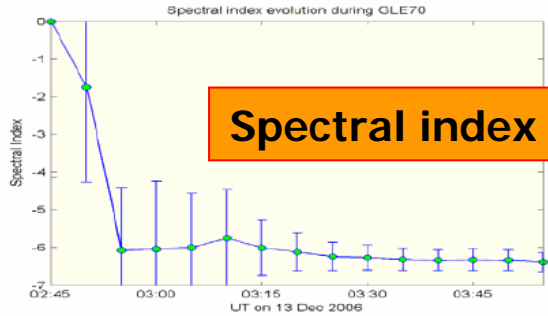
## Kernel of NM BANGLE





# GLE Modelling – NM BANGLE

## NM BANGLE - Results



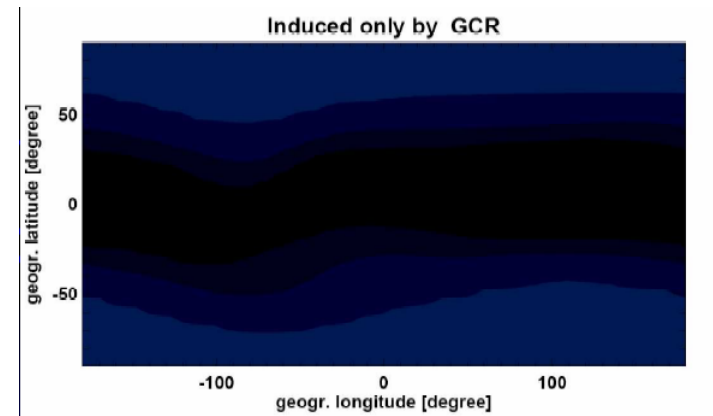
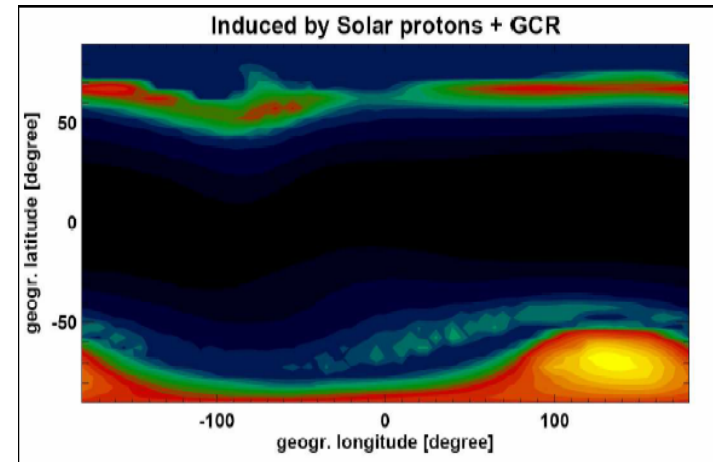
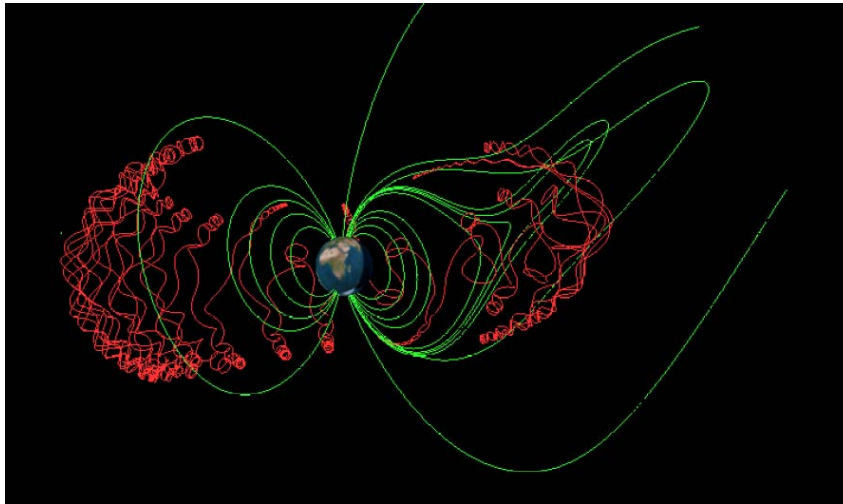




## Applications with 1-min real-time data

### Atmospheric Ionization – Planetocosmics code (UBERN)

The Planetocosmics code can illustrate particle trajectories affected by a magnetic field and calculate the ionization of the atmosphere during solar energetic particles







# Possible users of NMDB



Power Supply



Airway Co



Telecommunications

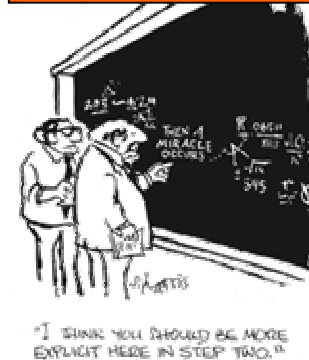


Space Agencies



Satellite Operation  
Centers

Scientists



Whoever interested



# NMDB Training Course, September 14-19, 2009, Athens, Greece

## Training course Facts & figures

- ✓ 55 students (PhD, Post-doc) from 16 countries
- ✓ 14 lectures & 4 lab courses
- ✓ 16 student presentations
- ✓ 1 Public talk

**Cosmic Rays and Neutron Monitors – A training course in science and applications**

September 14-19, 2009, Athens, Greece

**Topics**

1. Galactic & Solar Cosmic Rays: Origin and Propagation
2. The Heliosphere & its effects on cosmic rays
3. Cosmic ray measurements
4. Cosmic ray effects on Earth

**Lecturers**

- A. Sakhar (Ural)
- B. G. Mészáros (Hungary)
- C. Mészáros (Hungary)
- D. Simeonov (Bulgaria)
- E. Simeonov (Bulgaria)
- F. Simeonov (Bulgaria)
- G. Simeonov (Bulgaria)
- H. Simeonov (Bulgaria)
- I. Simeonov (Bulgaria)
- J. Simeonov (Bulgaria)
- K. Simeonov (Bulgaria)
- L. Simeonov (Bulgaria)
- M. Simeonov (Bulgaria)
- N. Simeonov (Bulgaria)
- O. Simeonov (Bulgaria)
- P. Simeonov (Bulgaria)
- Q. Simeonov (Bulgaria)
- R. Simeonov (Bulgaria)
- S. Simeonov (Bulgaria)
- T. Simeonov (Bulgaria)
- U. Simeonov (Bulgaria)
- V. Simeonov (Bulgaria)
- W. Simeonov (Bulgaria)
- X. Simeonov (Bulgaria)
- Y. Simeonov (Bulgaria)
- Z. Simeonov (Bulgaria)

**Public talk**

- A. Sakhar (Ural)

**Scientific Organizing Committee**

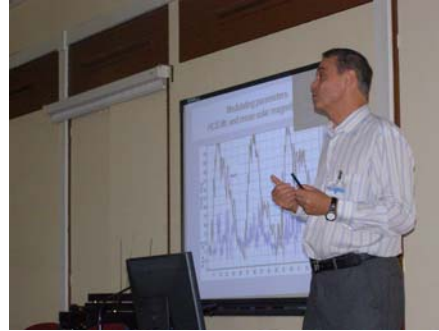
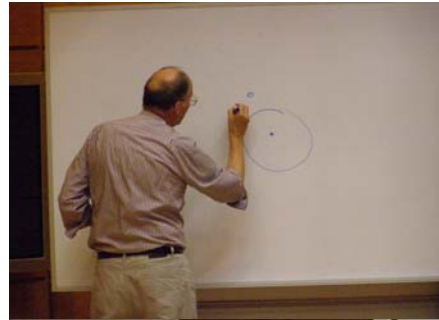
- E. Simeonov (Bulgaria)
- C. Mészáros (Hungary)
- A. Sakhar (Ural)
- F. Simeonov (Bulgaria)
- K. Simeonov (Bulgaria)

**Local Organizing Committee**

- B. G. Mészáros (Hungary)
- A. Sakhar (Ural)
- M. C. Papageorgiou (Athens)

**NMDB Collaborators**

More info: [www.nmdb.eu](http://www.nmdb.eu)  
Contact: [nmdb@phys.uoa.gr](mailto:nmdb@phys.uoa.gr)



✓ Visit at the Athens NM







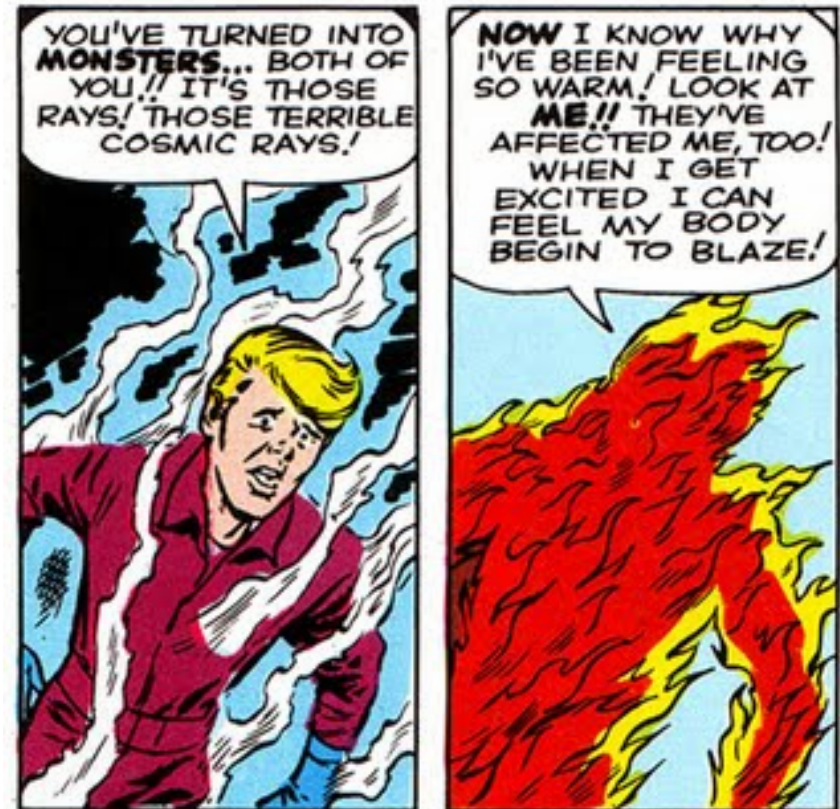
# Conclusions

- ◆ All NM stations have been upgraded providing high-resolution NM data in real-time
- ◆ High resolution NM data in real-time, are being collected and made public in common format
- ◆ A state of the art, centralized database with distributed mirrors, have been build
- ◆ An innovate registration system for cosmic rays have been implement
- ◆ Application tools using NM data have been developed
- ◆ A public outreach website with the scope to inform about cosmic rays and possible effects on humans, technological systems and environment have been created.
- ◆ The first ever Training Course on Cosmic Ray Physics and Applications have been realized

**NMDB database is an open tool  
All applications can be used as a service for scientists and for all**



# Fantastic 4 Effect !





Thank you for your attention 😊



Project's website:

<http://www.nmdb.eu>

